

# BROAD BAND SEDS OF HIGH-Z RADIO GALAXIES

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**4C 23.56**

Restframe Wavelength ( $\mu\text{m}$ )

$z=2.46$

$10^{-4} \ 10^{-3} \ 0.01 \ 0.1 \ 1 \ 10 \ 100 \ 1000 \ 10^4 \ 10^5 \ 10^6$

*Miley G. & De Breuck C., 2008*

$\nu f_\nu \ (10^{-15} \ \text{erg s}^{-1} \ \text{cm}^{-2})$

100

100

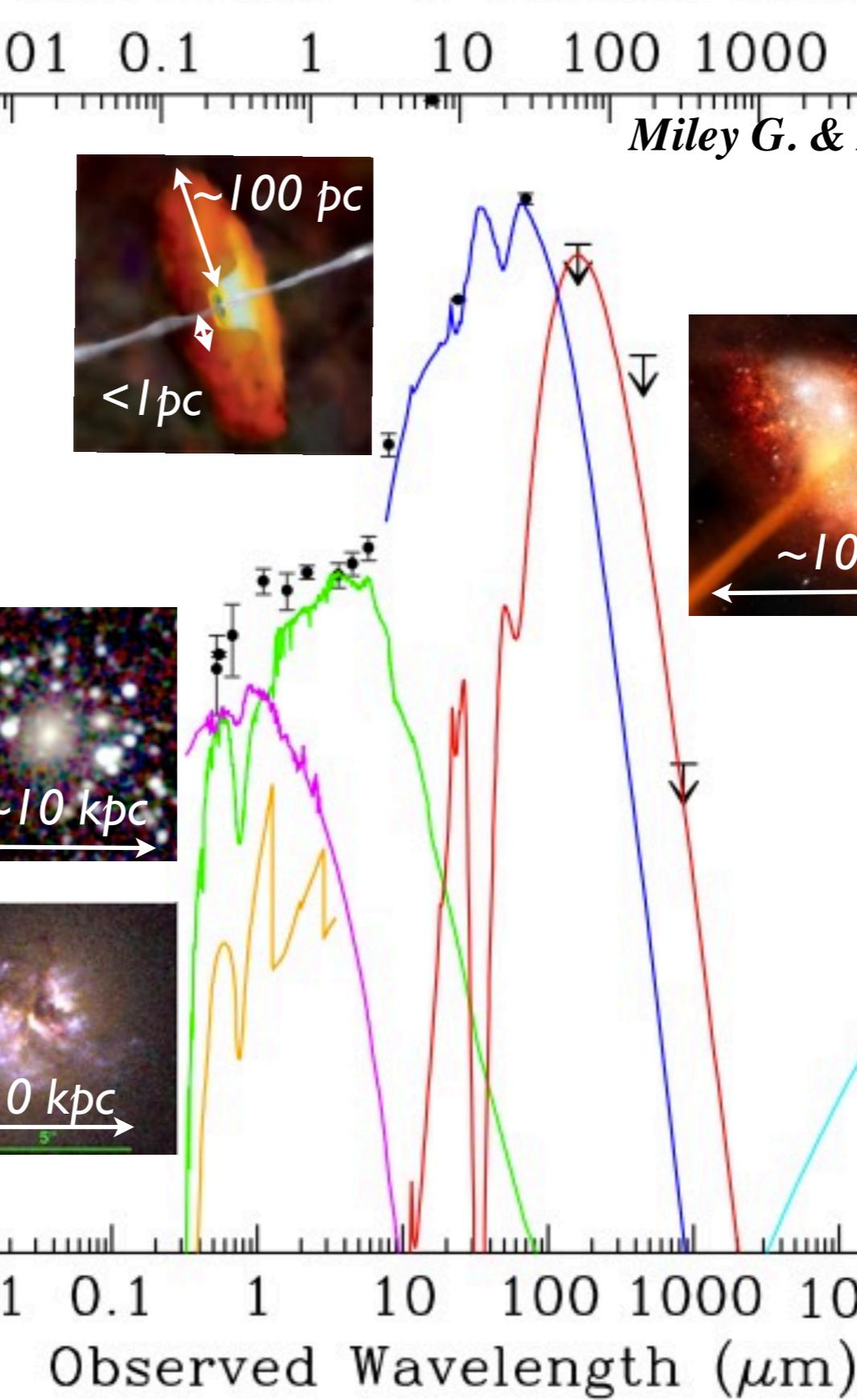
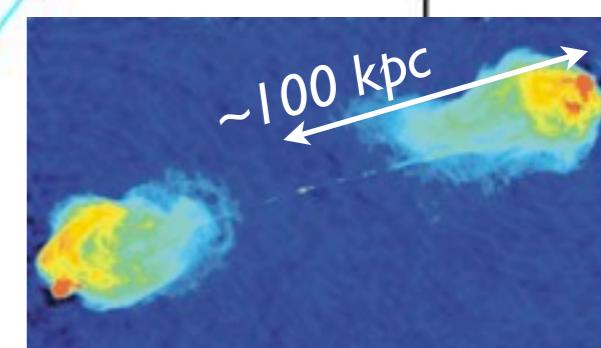
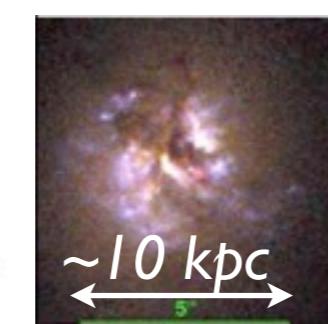
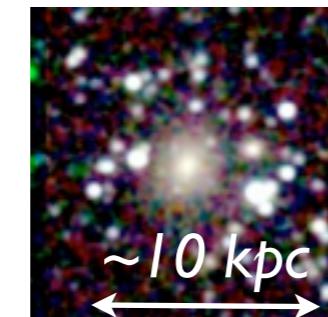
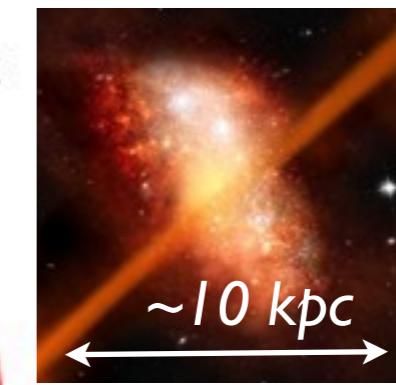
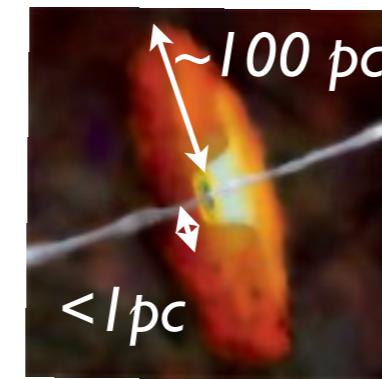
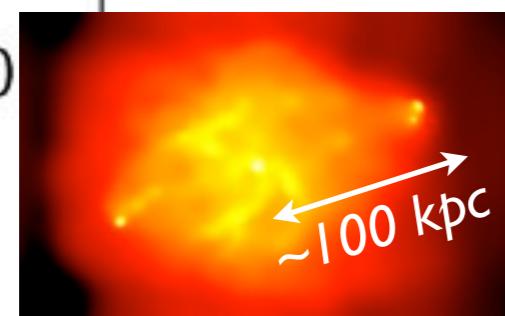
10

1

X-rays  
neb. cont.  
scat. AGN  
stars  
hot dust  
cold dust  
synchrotron

$10^{-4} \ 10^{-3} \ 0.01 \ 0.1 \ 1 \ 10 \ 100 \ 1000 \ 10^4 \ 10^5 \ 10^6 \ 10^7$

Observed Wavelength ( $\mu\text{m}$ )





# PROJET HERGÉ

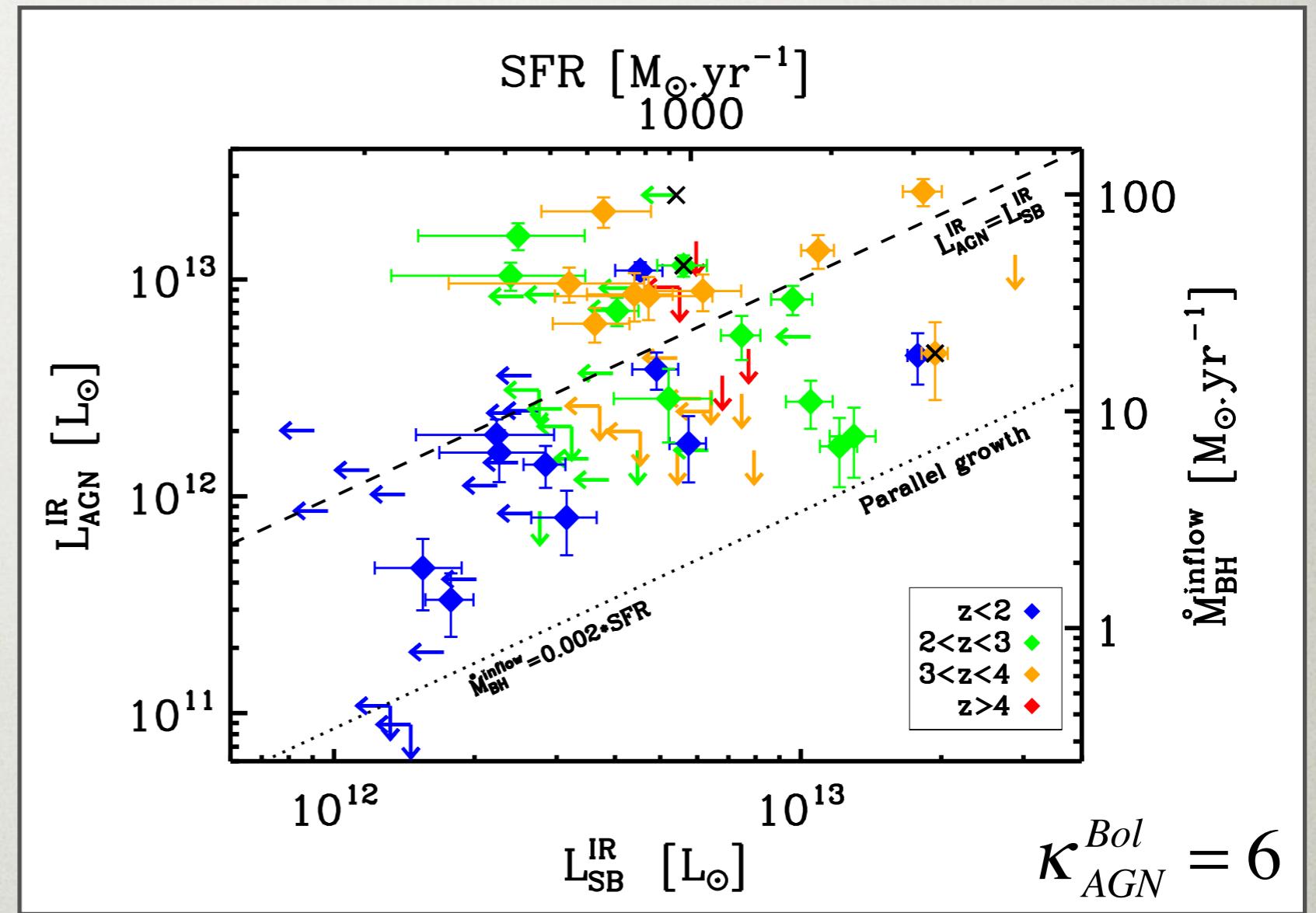
*Herschel* Radio Galaxy Evolution project  
International team (~25 members)

- 70 powerful radio galaxies at  $1 < z < 5.2$ : Mostly USS radio source selected at 3GHz restframe, parent sample >200 objects, all-sky
- Supporting data: *HST*, VLT, Keck, *Spitzer*, *Herschel*, SCUBA, LABOCA, VLA, ATCA, MWA, ALMA
- SED coverage from UV to radio (>20 broad bands)

*Some publications:* Seymour et al. 2007, Miley & De Breuck 2008, De Breuck 2010, Galametz et al. 2012, Mayo et al. 2012, Seymour et al. 2012, Ivison et al. 2012, Wylezalek et al. 2013, Rocca et al. 2013, Drouart et al. 2014, Emonts et al. 2015-2016, Gullberg et al. 2016, Drouart et al. 2016, T. Falkendal in prep.

# WHAT'S OLD(ER)?

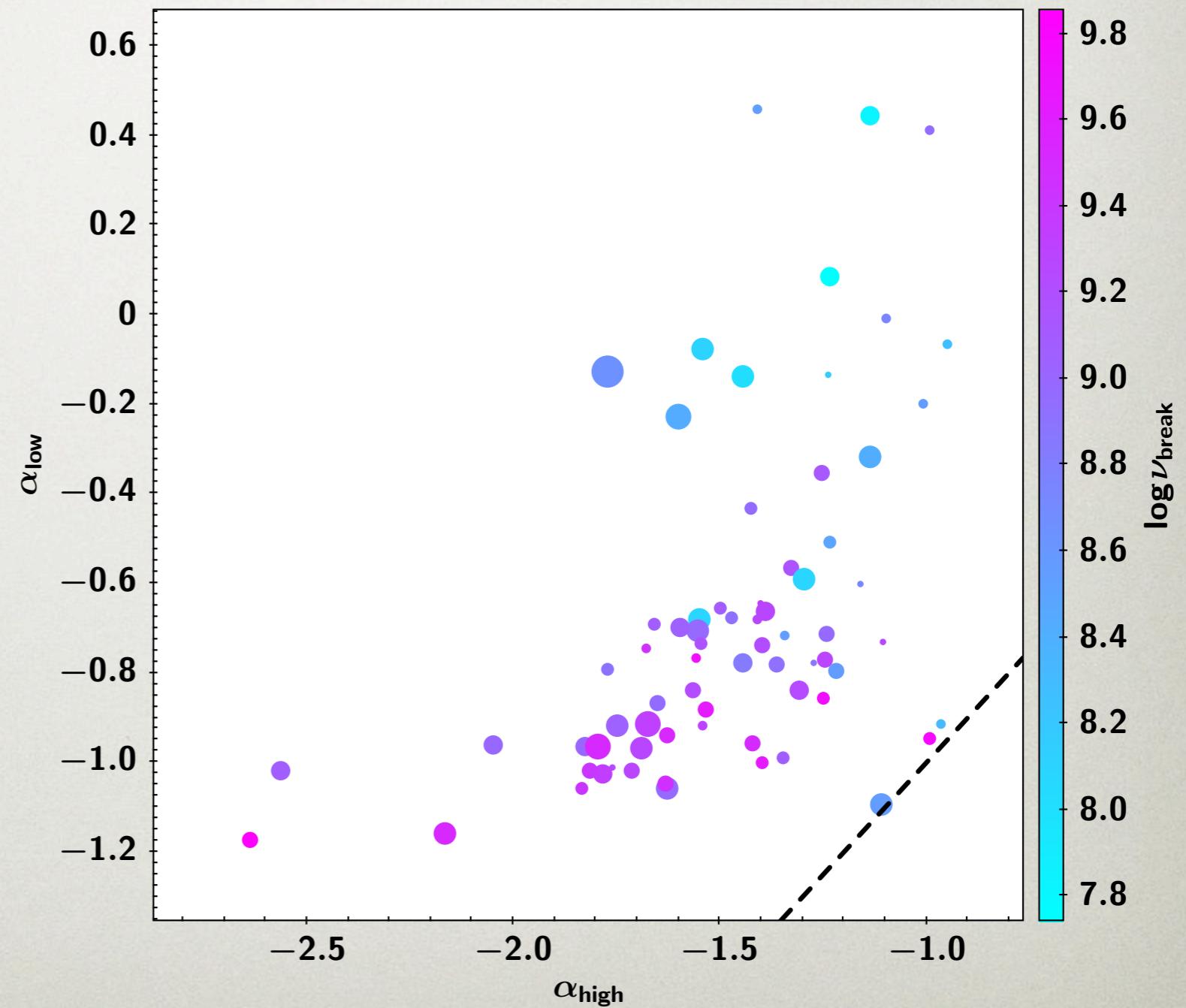
- HzRGs are massive galaxies sitting in denser environments (Seymour et al. 2007, De Breuck et al. 2010, Wylezalek et al. 2013,...)
- The radio emission betrays the presence of a SMBH having an impact on its environment (e.g. Nesvadba et al. 2009)
- They are the progenitors of the massive ellipticals in the local Universe (Rocca et al. 2004, Labbe et al. 2005)
- they are type 2 AGN (Obscured): Excellent laboratories to study host, to understand feedback impact and SMBH-host relation (e.g. Nesvadba et al. 2016, Drouart et al. 2016)



Drouart et al. 2014

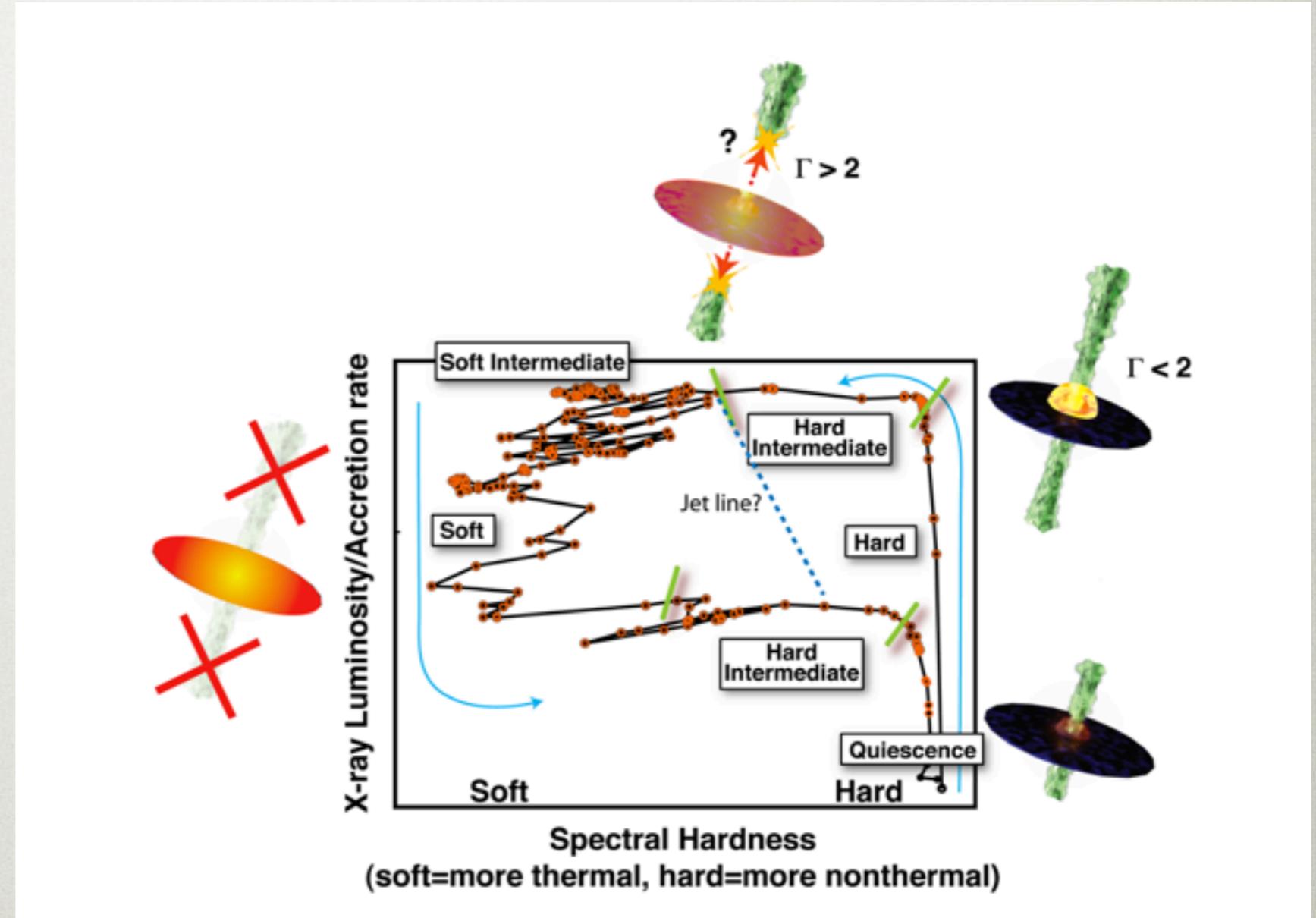
# WHAT'S NEW(ER)?

- Fitting a double power law with a break frequency (MrMoose)
- At least 20 datapoints on most of the sample, covering the 50MHz-20GHz range
- Size refers to redshift
- Two different type of sources as the break do not correspond to the same physical process
- Contains environmental and intrinsic property information



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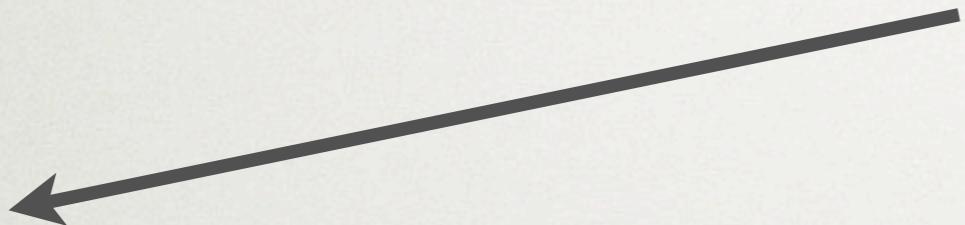
- Local galactic black holes and supermassive black hole are similar and some properties are scale invariant (Plotkin et al. 2015)
- The amount of energy emitted mechanically (jet) and radiatively (quasar) are similar (Nesvadba et al. 2017)
- X-rays data are very useful to assess this diagnostic as less sensitive to dust obscuration
- High synergy with radio required to create the full picture => multi-wavelength dataset



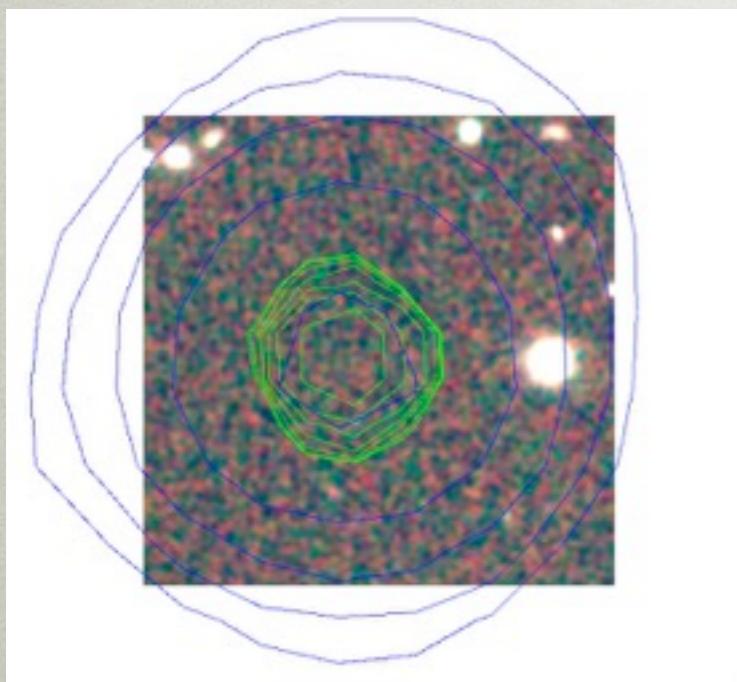
# WHAT'S FOR?

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better SEDs of HzRGs in the  $1 < z < 5$  range



Templates to push redshift  
limit,  $z > 6$



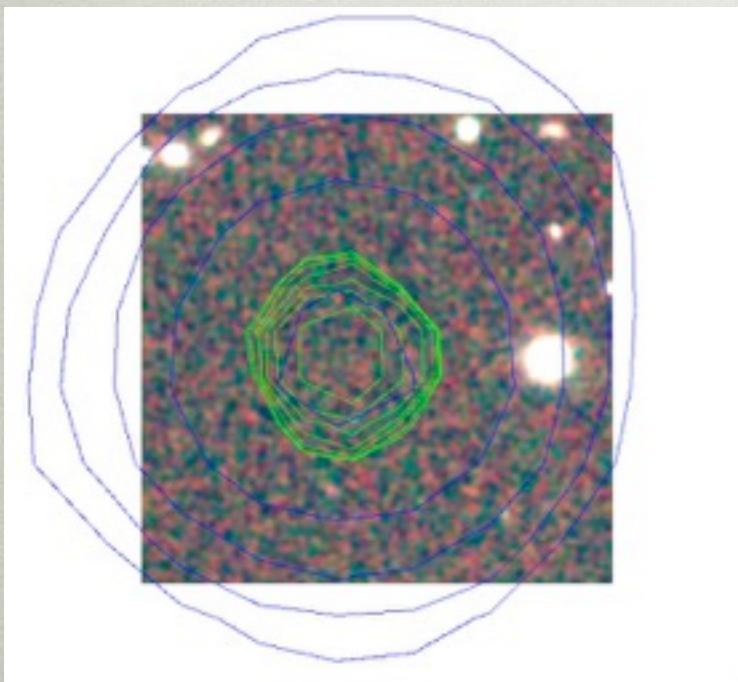
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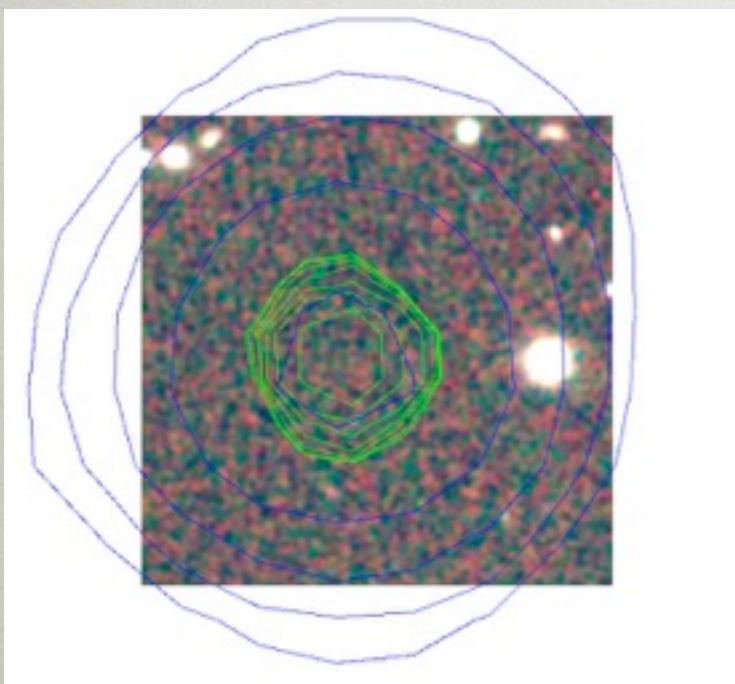
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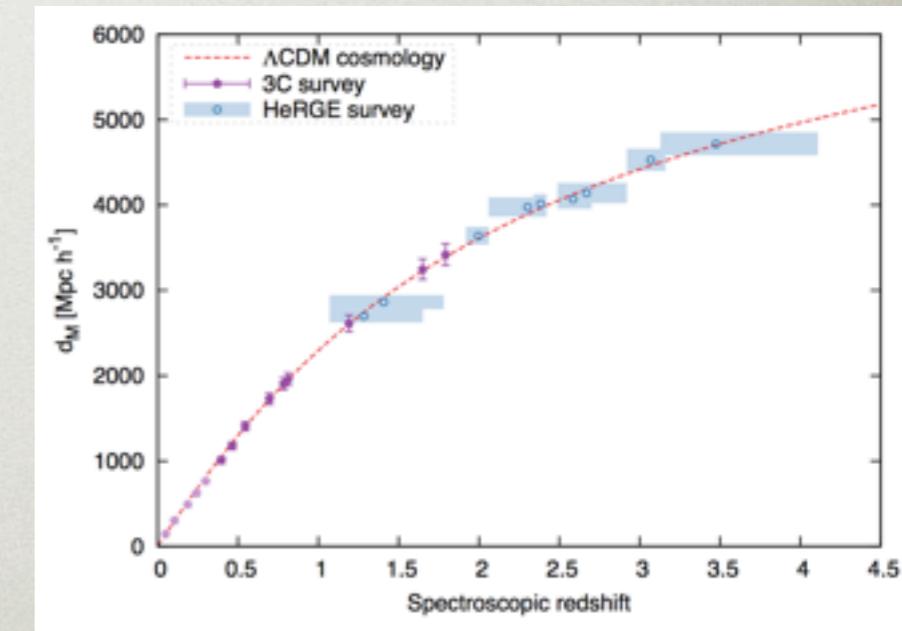
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feed this knowledge for redshift estimates in radio



Turner et al., sub.



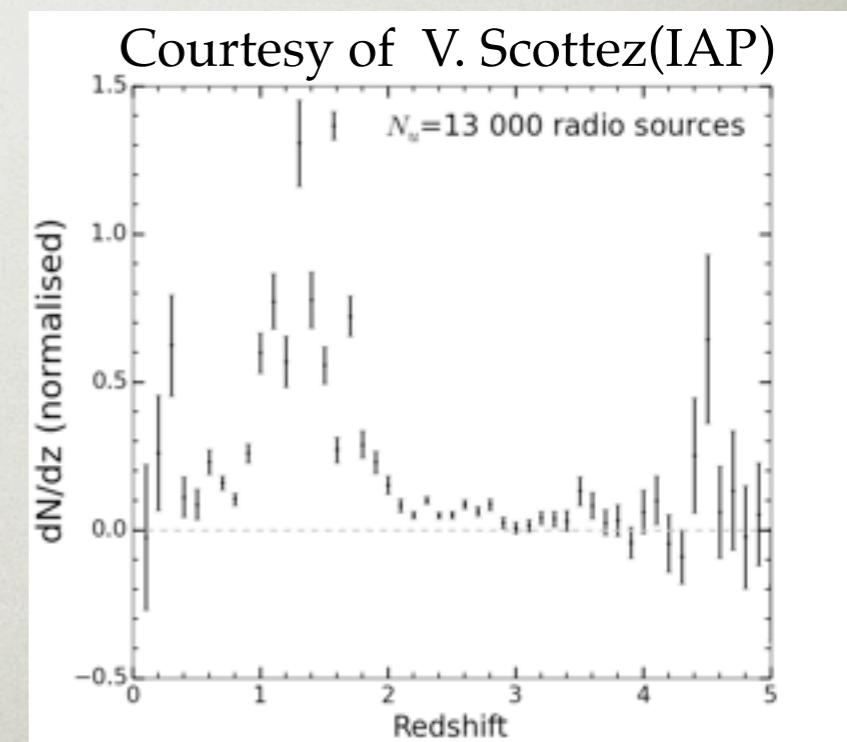
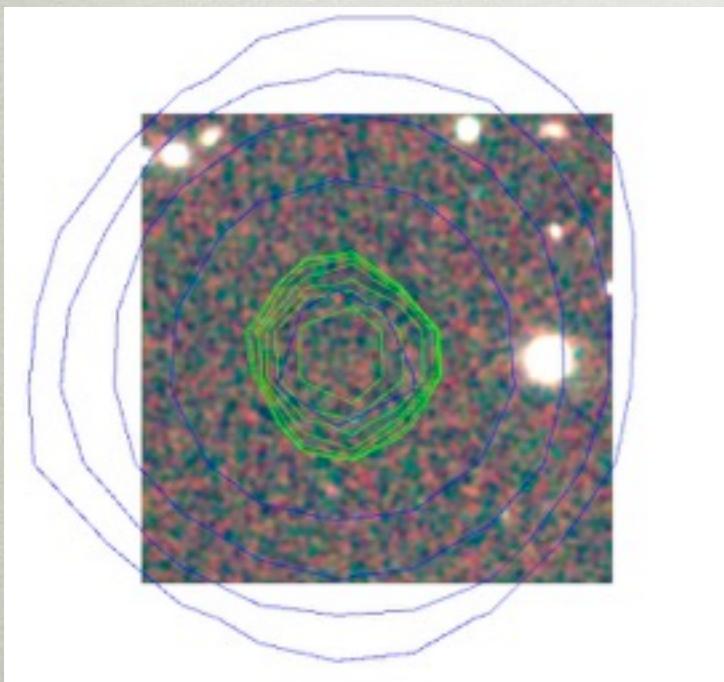
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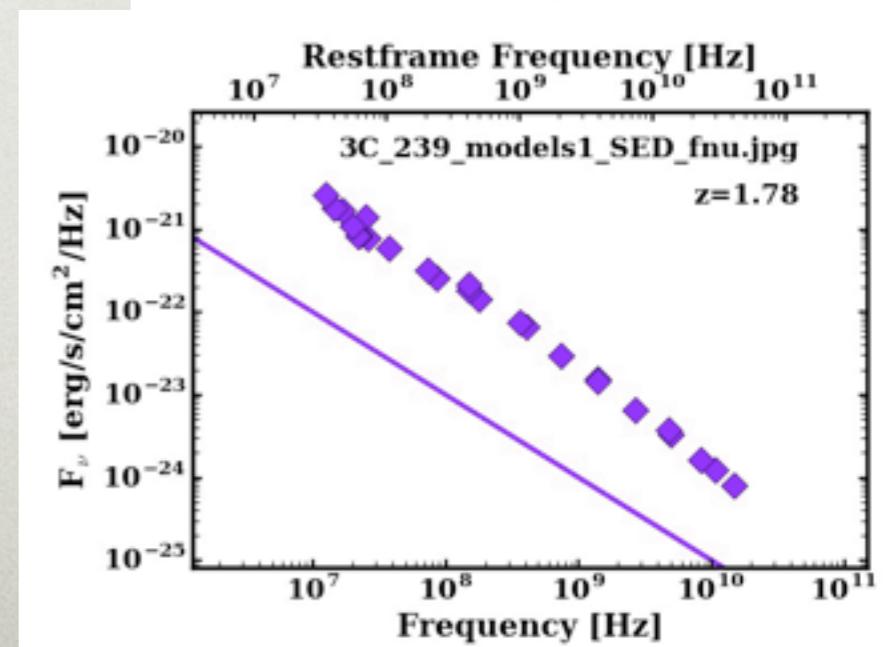
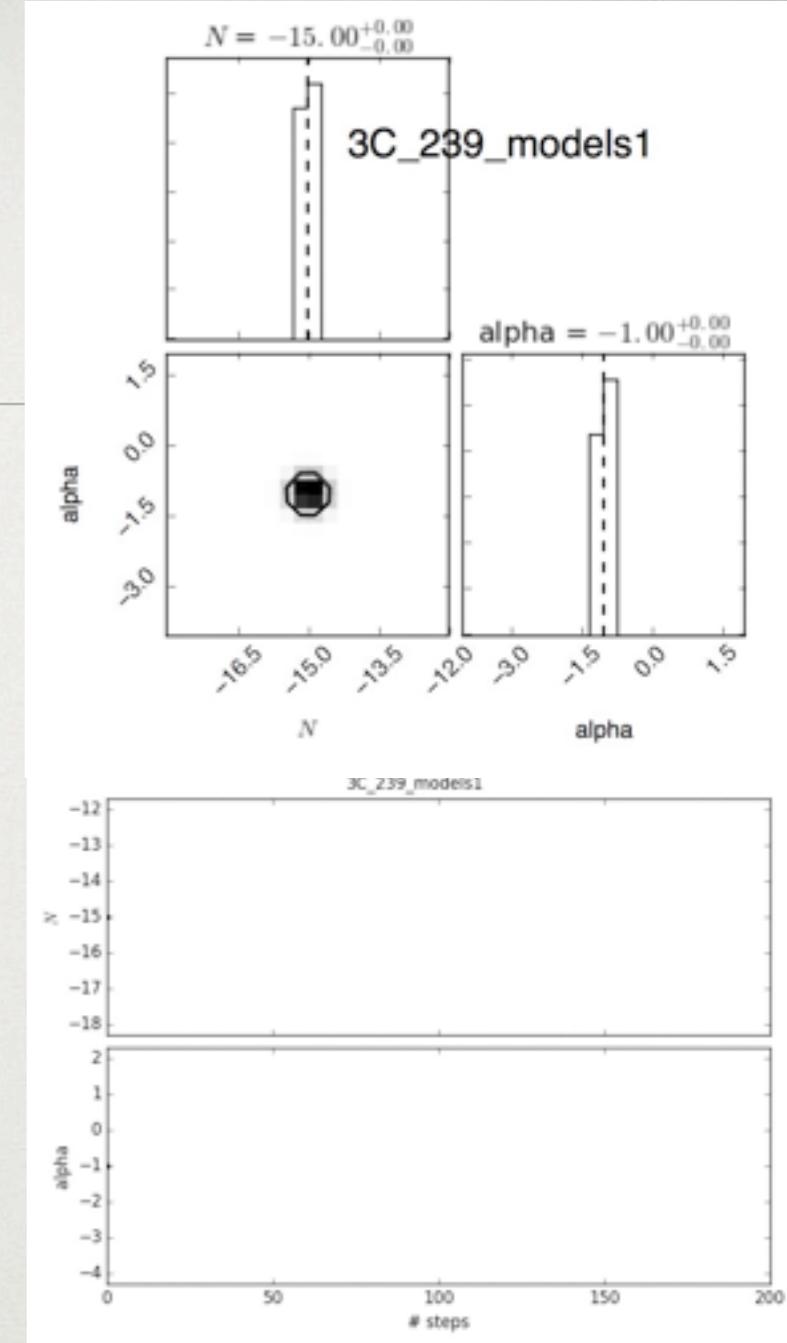
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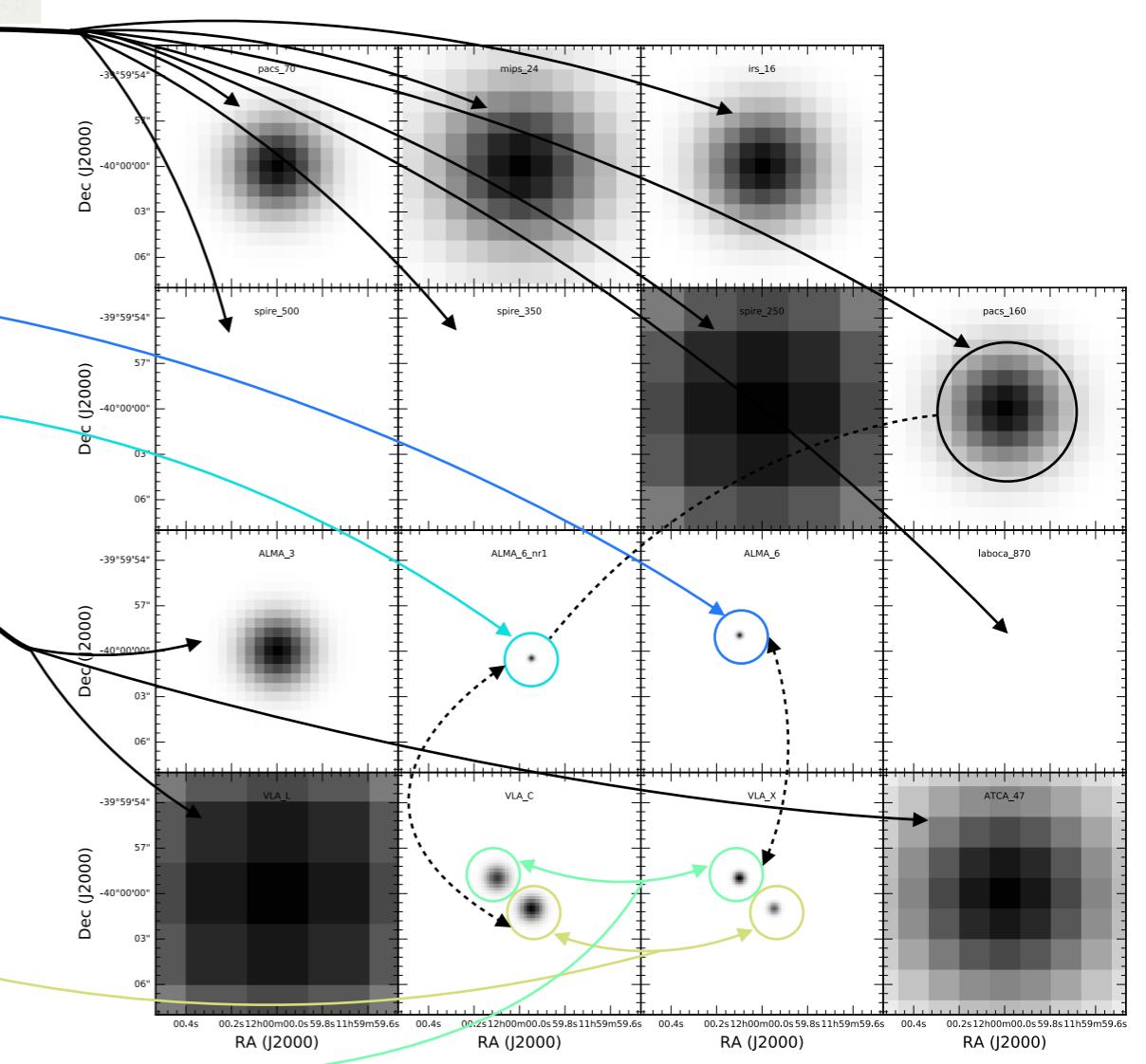
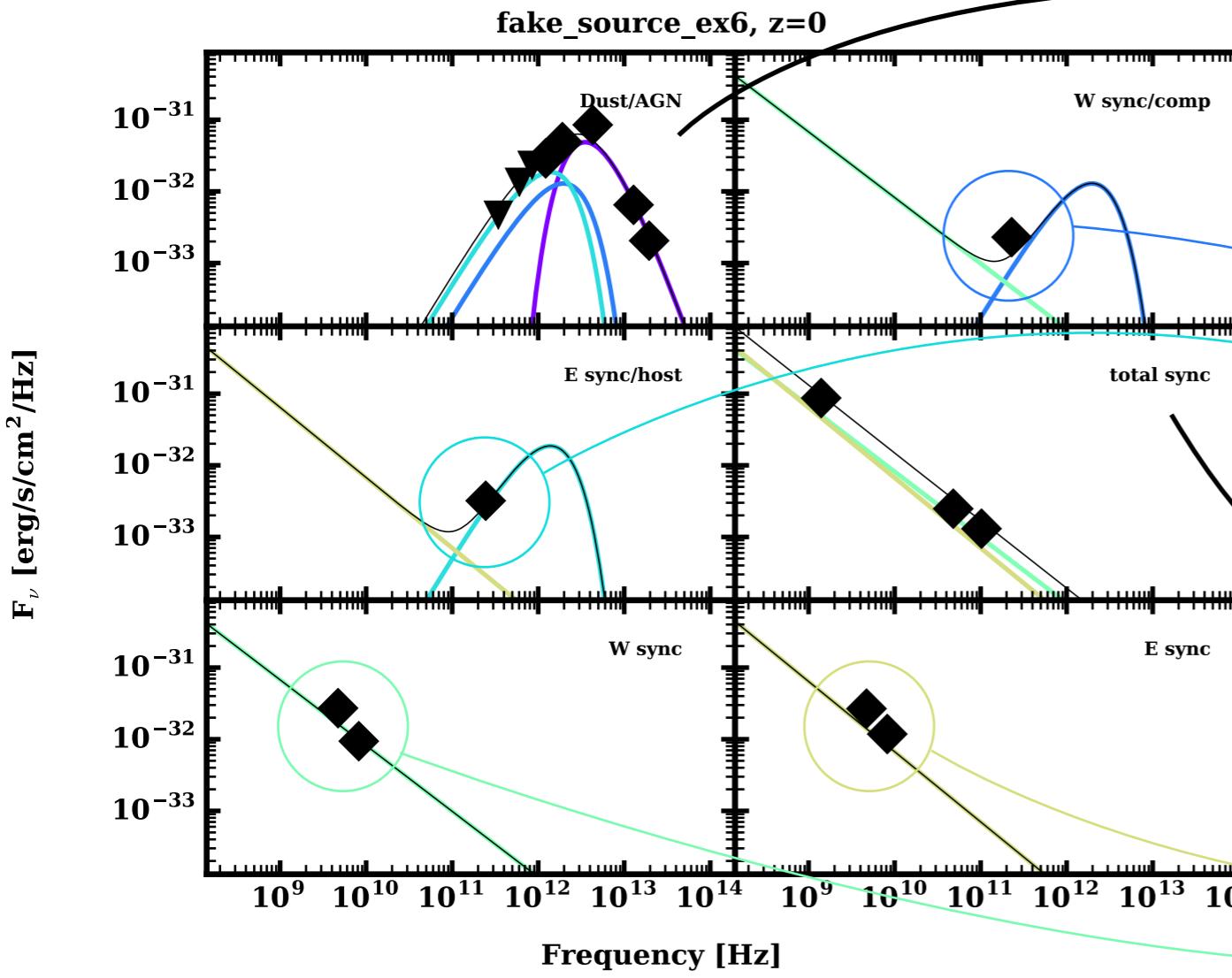
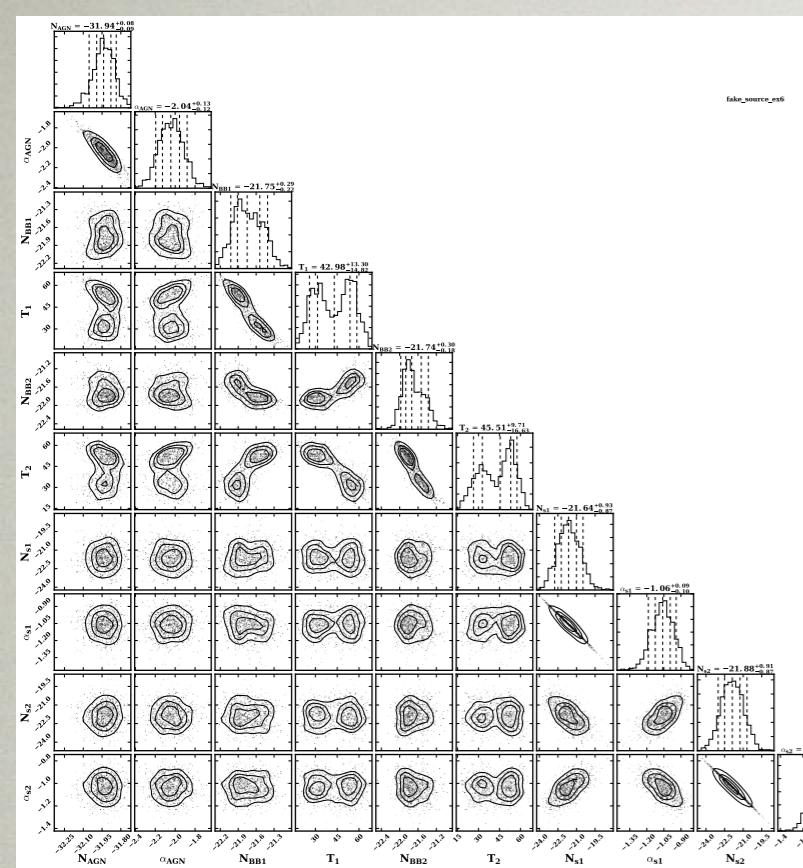
# A WORD ABOUT SED FITTING

# MRMoose (ANOTHER) FITTING TOOL

- Bayesian framework
- handle upper limits properly
- handle blended source
- user-managed model / filter database
- User-friendly not user-opaque
- some limitations for v1.0 (public release, next week)
- Python 2.7 (hopefully 3.0 in close future)



# MRMOOSE - A REALISTIC CASE



# CONCLUSIONS

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- Low Frequency radio astronomy to open a new window for AGN and star forming galaxy physics / demographics
- Synergy essential with the high frequency side to obtain the full picture
- SMBH in HzRG: similar output energy in total radio and IR but over different timescale! Is the radio phase the prelude to the quasar phase?
- MrMoose: a great tool to the multi-wavelength astronomer!